

20171020 STATUS REPORT (2017 Fall KPS presentation)

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TOF description

Geometry

An array of plastic scintillation counters.

One bar = $10 \times 5 \times 170 \text{ cm}^3$

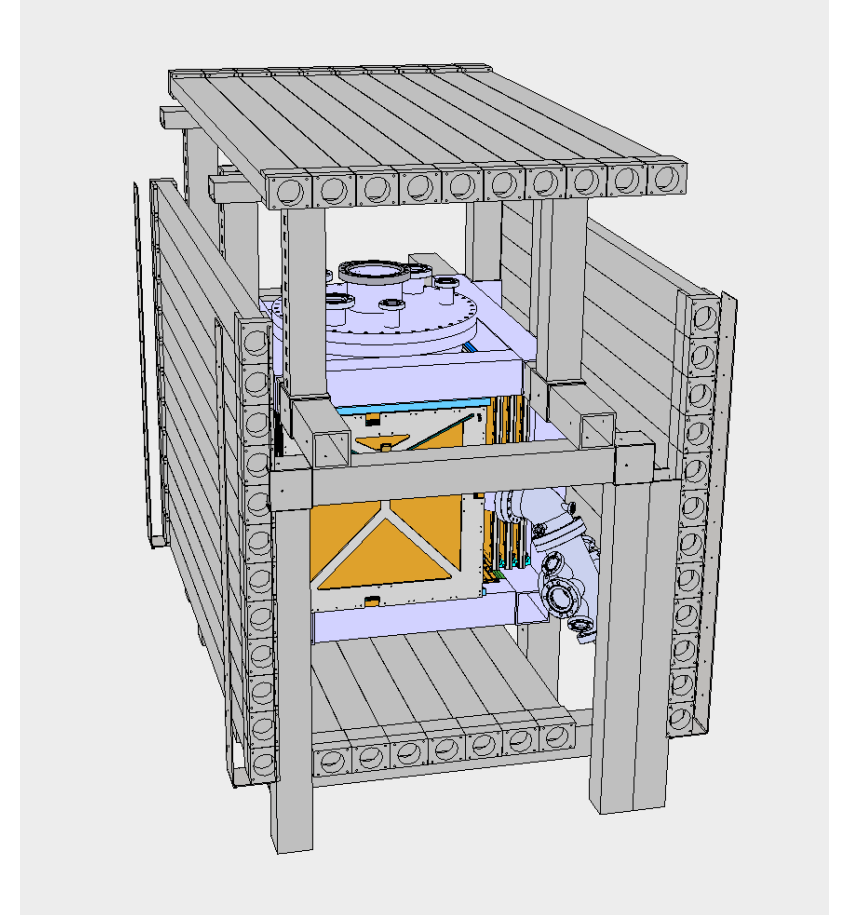
One counter = (PMT) – (plastic bar) – (PMT)

One wall (left/right) = 12 counters

(top/bottom) = 10 counters

* It is changeable, depending on FFC

Totally, 44bars and 88PMTs



TOF description

Purpose

From the fastest one among signals from the annihilation,
Free-fall time is obtained.

From the time difference between top & bottom hits,
top annihilation, bottom annihilation,
and cosmic ray hit are distinguished.

$dT < 0$ for chamber-bottom-annihilation
 $0 < dT < 2$ top
 $2 < dT$ for cosmic-ray signal

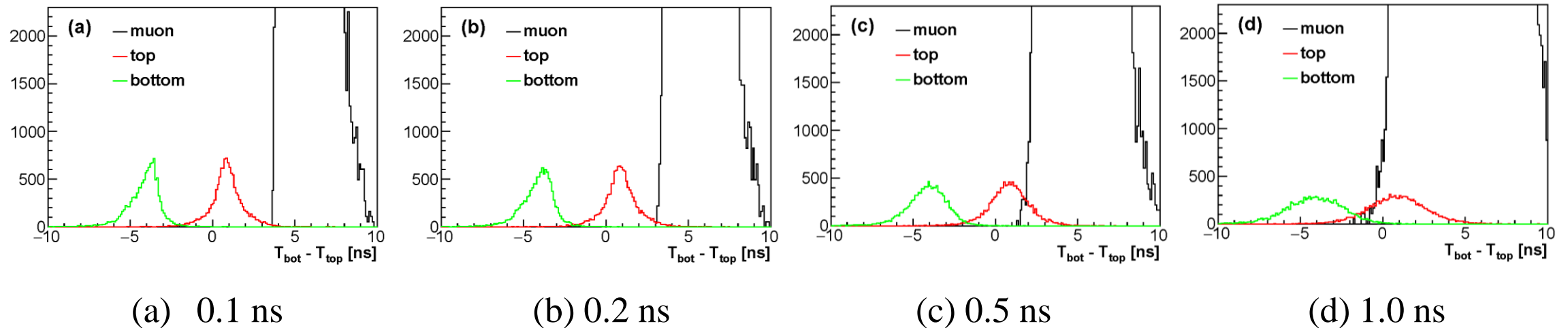
T_1 = time to bottom of TOF
 T_2 = time to top of TOF
 $dT = T_1 - T_2$

TOF description

Requirement

To cosmic ray rejection, the time resolution should be smaller than 0.2ns.

3-pi decay events, 825mm height, 4 different t-resolution
(normalized pion:muon = 1:5500)



Distinguishable !

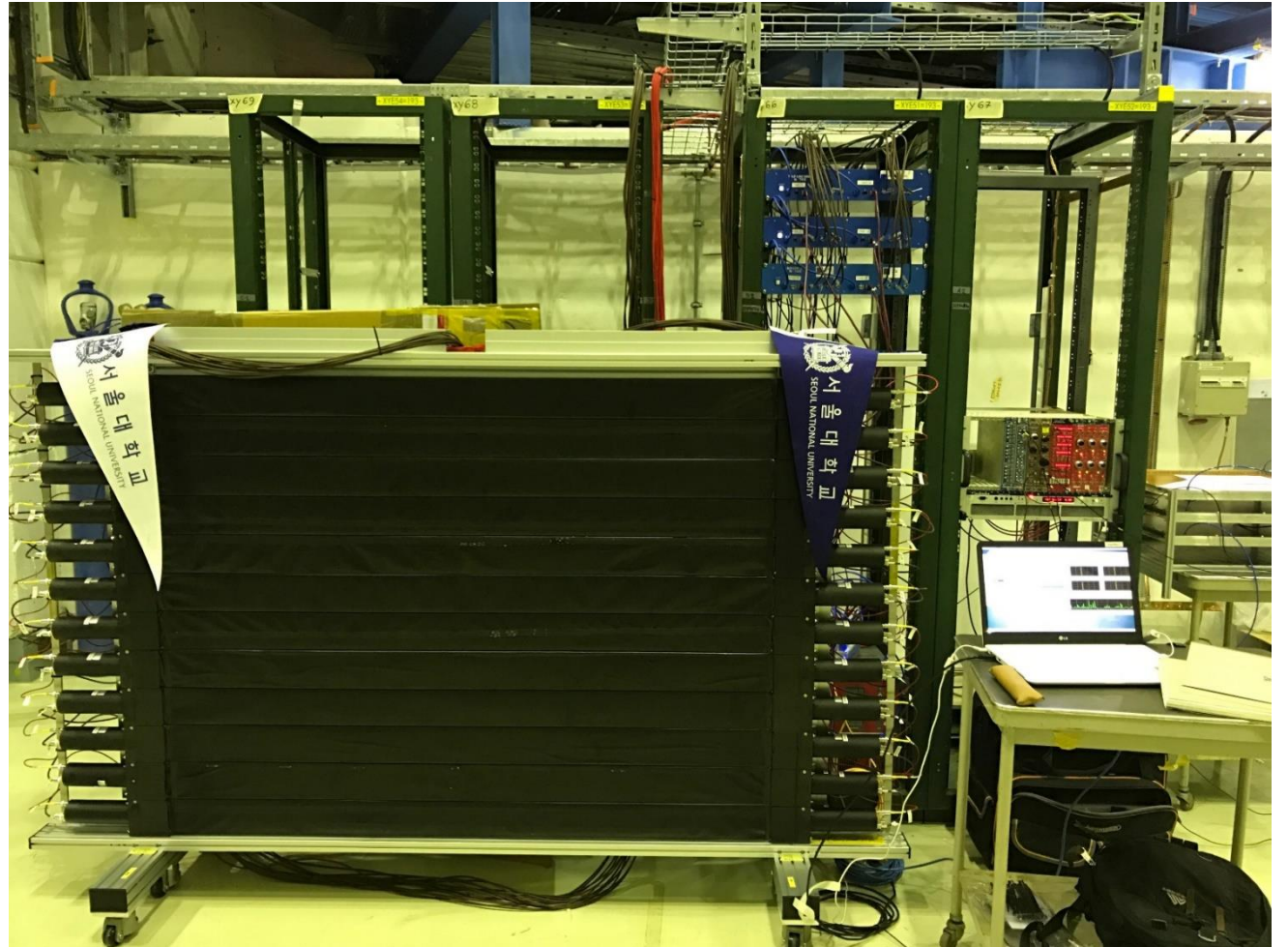
Chamber-top-annihilation signal is not distinguishable from cosmic ray signal

Performance

Cosmic muon test : settings

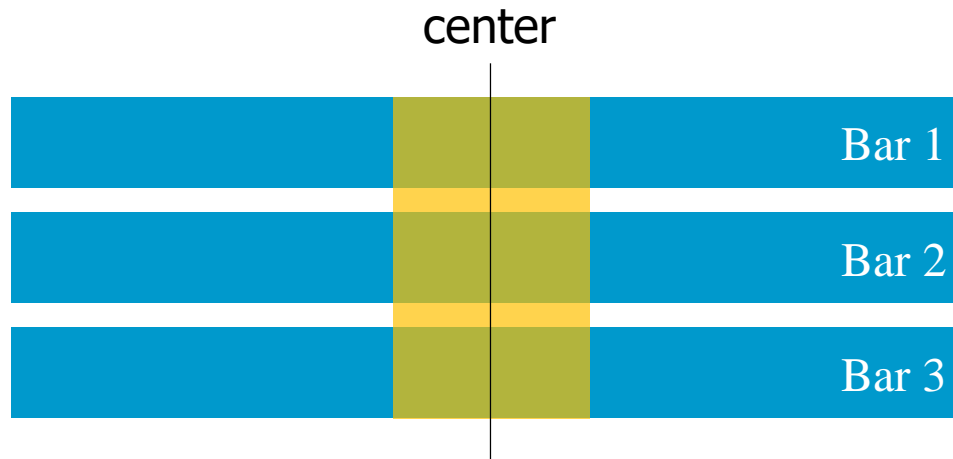
A wall of TOF has been installed in AD hall, CERN, during this summer.

To check DAQ program, and to see its performances as a wall, Cosmic muon test has been proceeded.



Performance

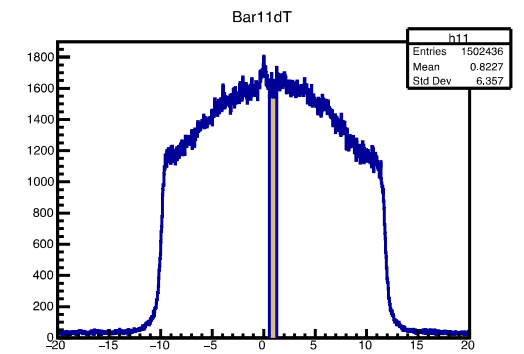
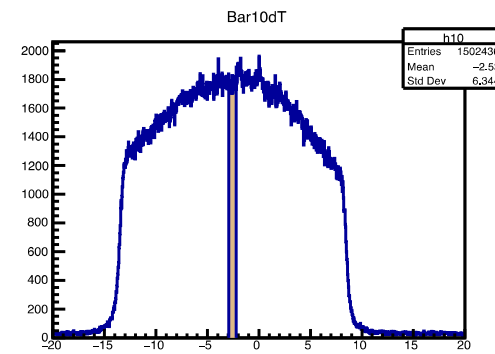
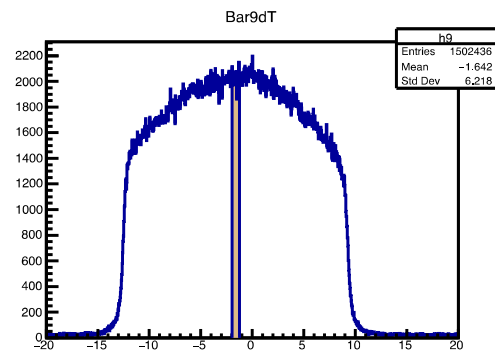
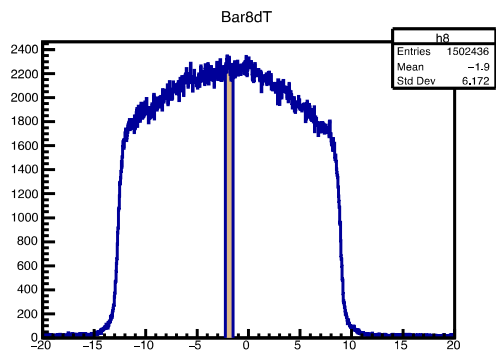
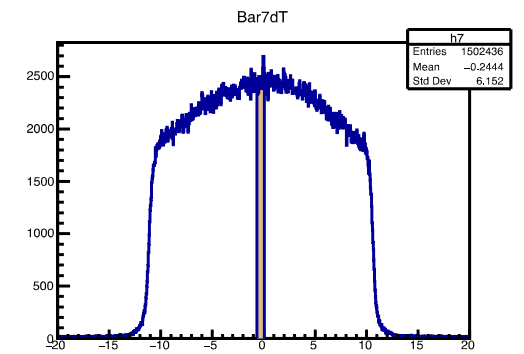
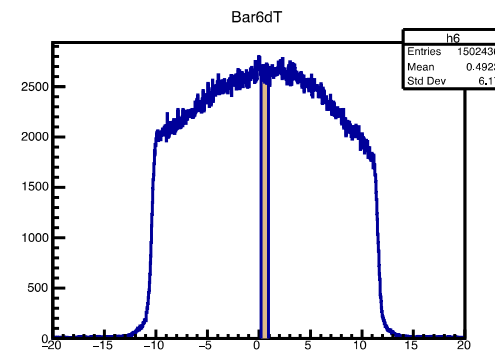
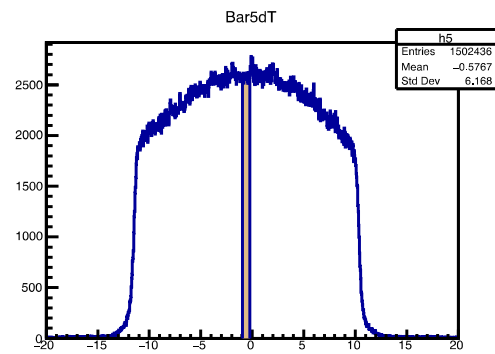
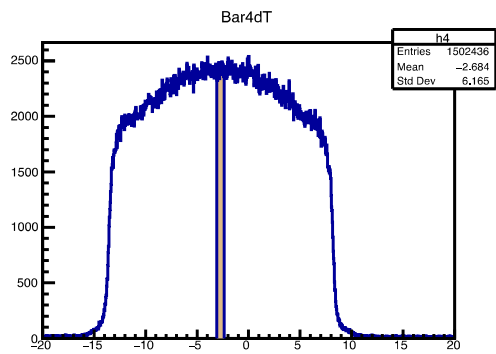
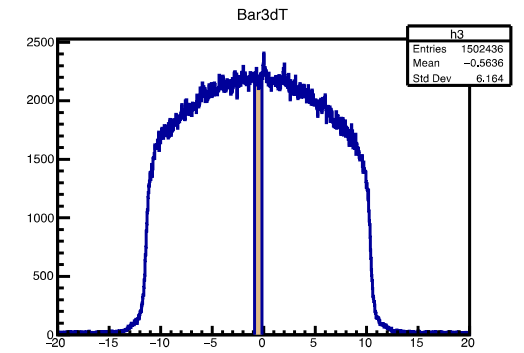
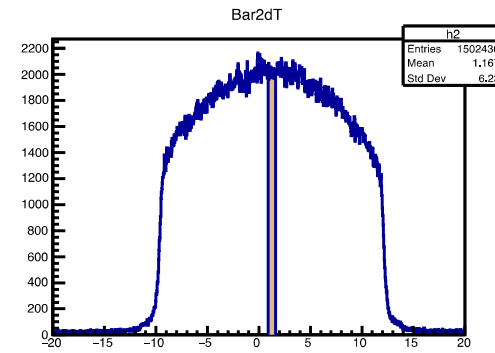
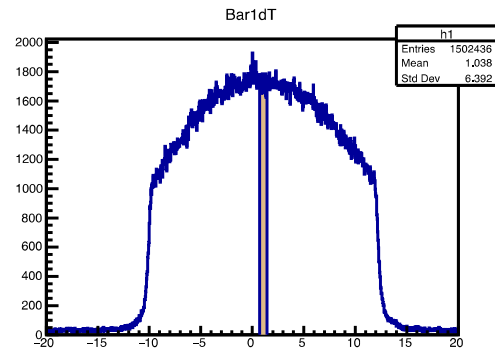
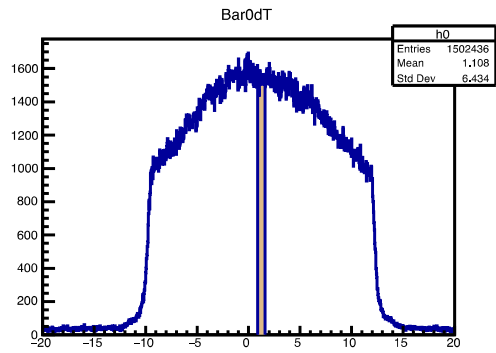
Time-resolution test



The event that cosmic ray hits any 6 bars among 12 bars, is saved.

Group three bars into a group.

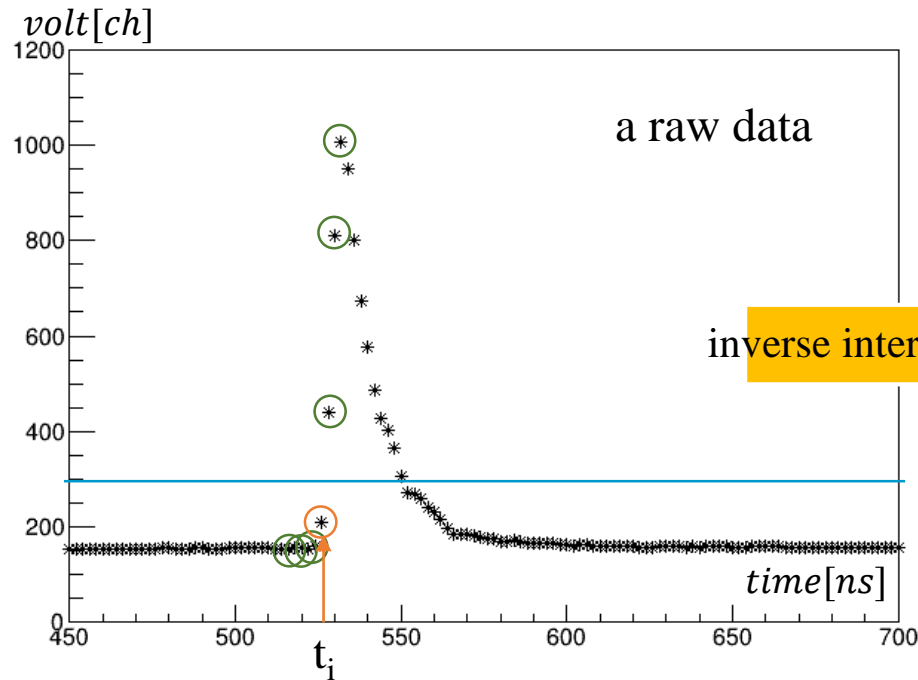
In a group,
select the events that passing through
within $\pm 5\text{cm}$ from their center(0cm).



Performance

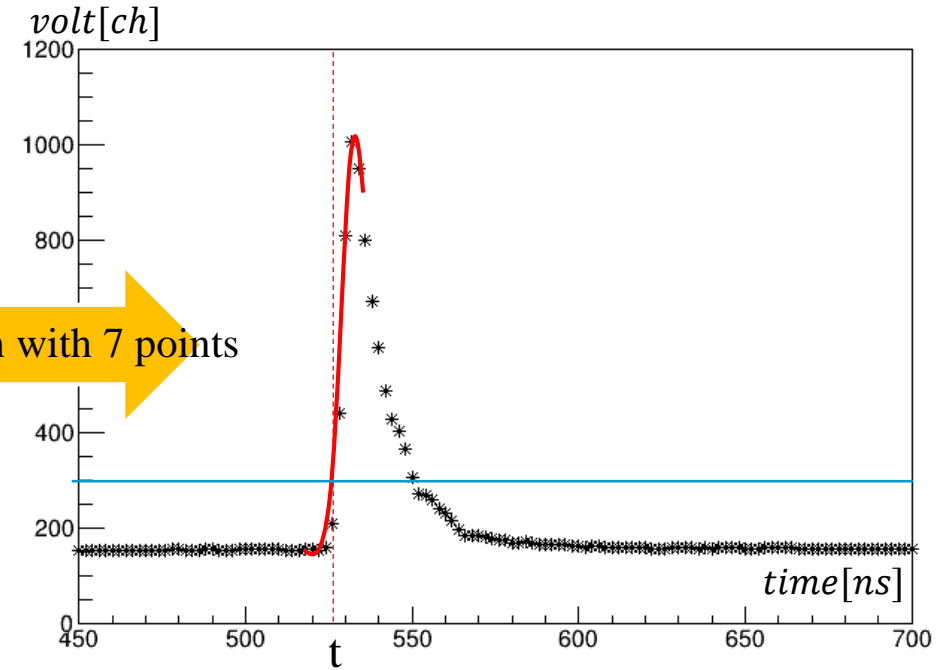
Time-resolution test

The event time of signal is determined by inverse interpolation method.



t_i = the nearest point from 10% line

inverse interpolation with 7 points

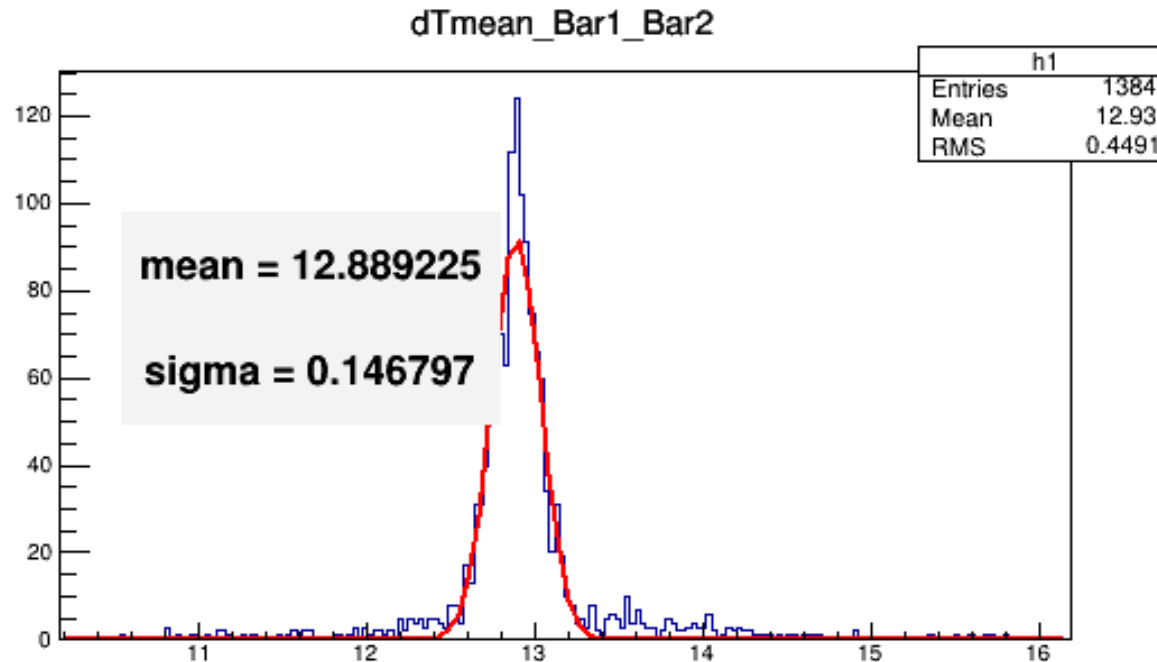


$$f(t) = 0.1 * (\text{peak height})$$

Performance

Time-resolution test

From the distribution of the event time difference between bars,
Time resolution of one bar can be calculated.

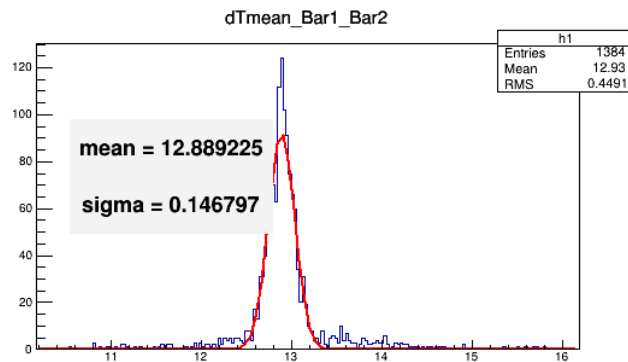


$$\begin{aligned} dtmean &= (tmean_1) - (tmean_2) \\ tmean_1 &= (t_1 + t_2) / 2 \\ tmean_2 &= (t_3 + t_4) / 2 \end{aligned}$$

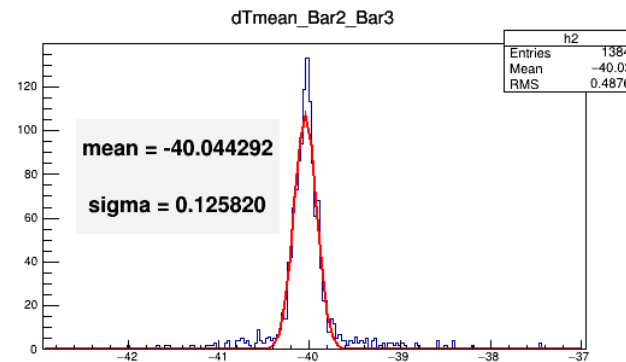
Performance

Time-resolution test

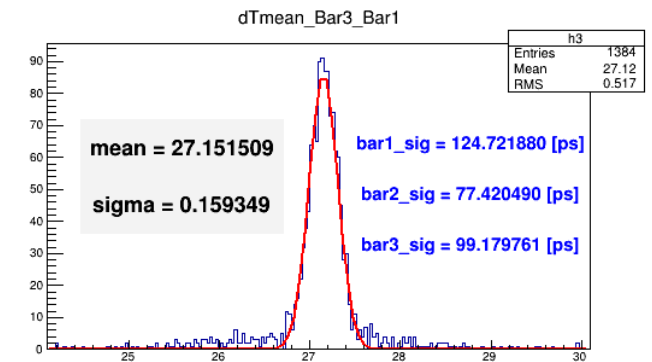
From the distribution of the event time difference between bars,
Time resolution of one bar can be calculated.



$$(\sigma_{h1})^2 = \sigma_1^2 + \sigma_2^2,$$



$$(\sigma_{h2})^2 = \sigma_2^2 + \sigma_3^2,$$

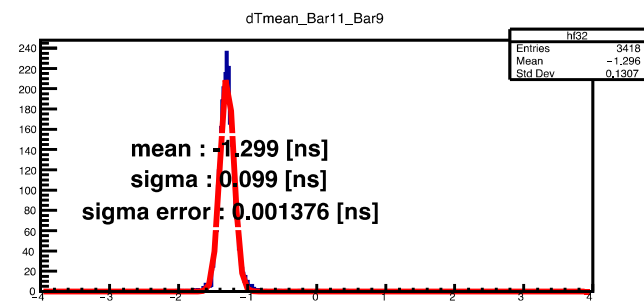
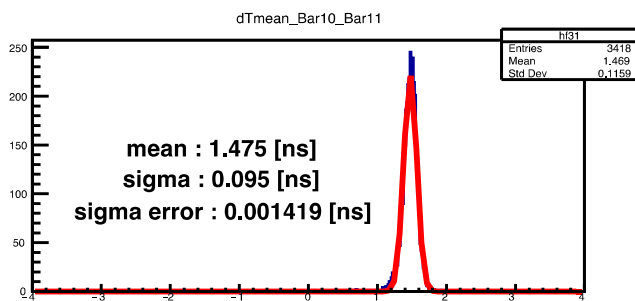
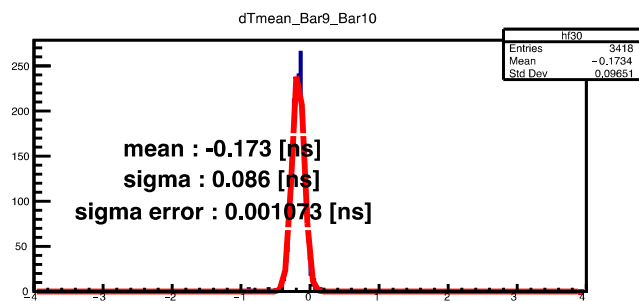
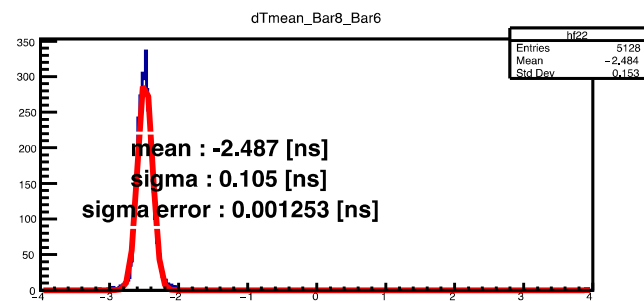
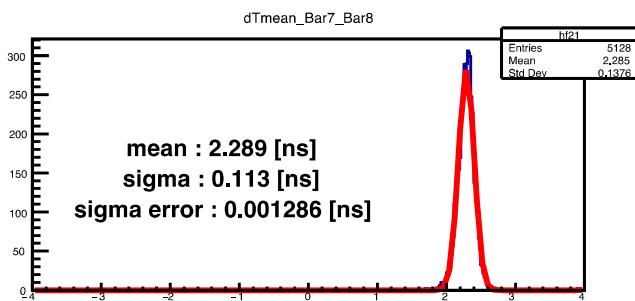
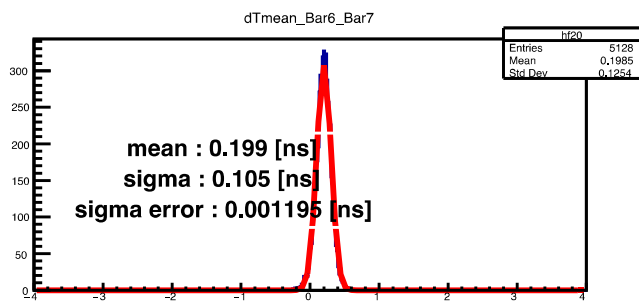
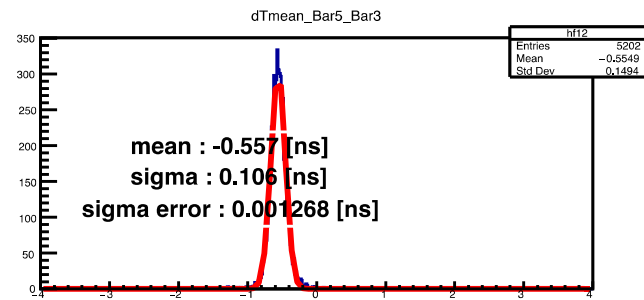
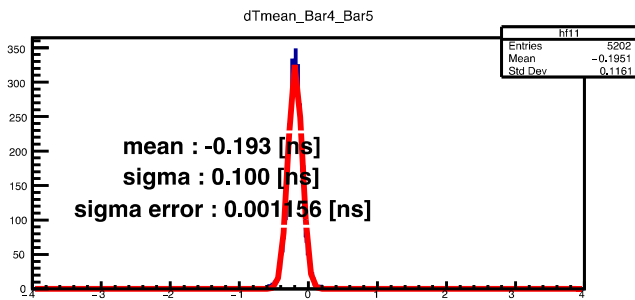
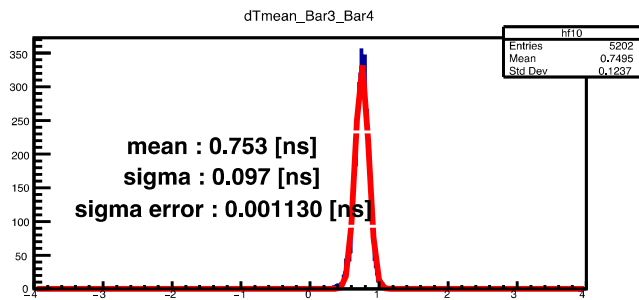
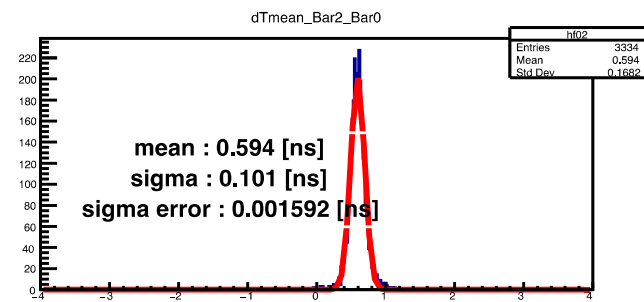
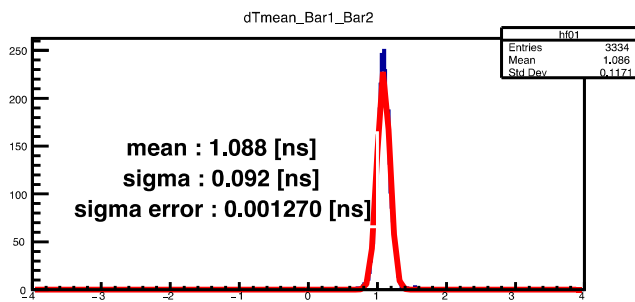
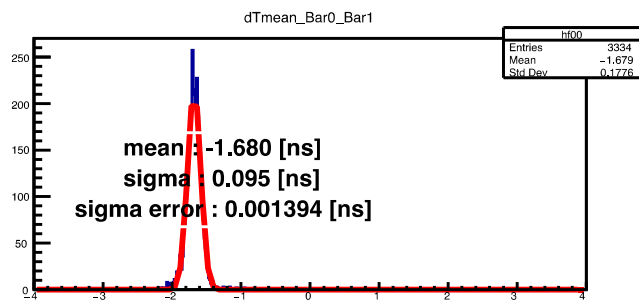


$$(\sigma_{h3})^2 = \sigma_3^2 + \sigma_1^2$$

$$\longrightarrow \sigma_1 = \sqrt{\frac{(\sigma_{h1})^2 - (\sigma_{h2})^2 + (\sigma_{h3})^2}{2}},$$

$$\sigma_2 = \sqrt{\frac{(\sigma_{h1})^2 + (\sigma_{h2})^2 - (\sigma_{h3})^2}{2}},$$

$$\sigma_3 = \sqrt{\frac{-(\sigma_{h1})^2 + (\sigma_{h2})^2 + (\sigma_{h3})^2}{2}}$$



Performance

Time-resolution test : results

bar#	Resolution[ps]		bar#	Resolution[ps]	
	prototype	A wall		prototype	A wall
1	96	73	7	78	68
2	87	60	8	98	79
3	124	69	9	96	80
4	77	73	10	83	63
5	99	64	11	83	58
6	100	77	12	81	76

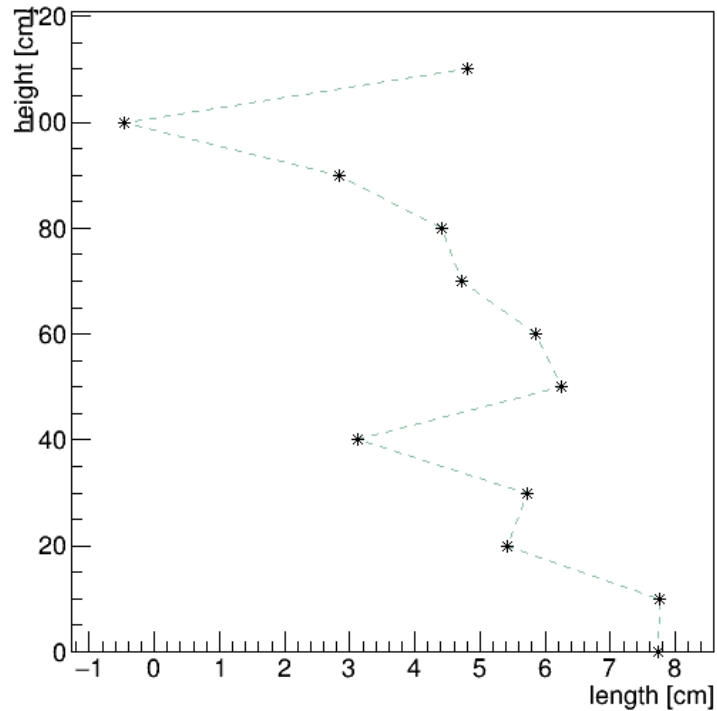
$$\sigma_{time} < 0.2 \text{ ns},$$

→ Enough to distinguish **top, bottom annihilations and cosmic ray signal**

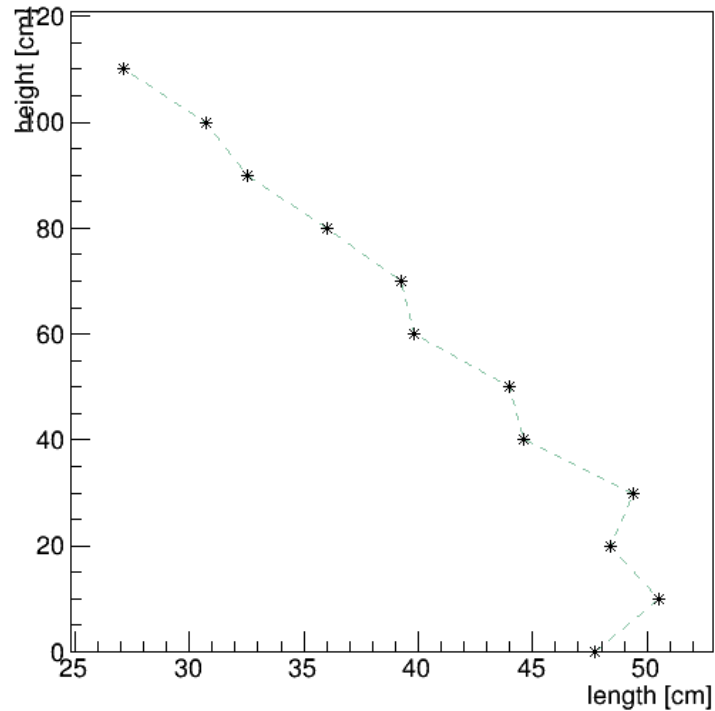
Performance

2D trajectory reconstruction

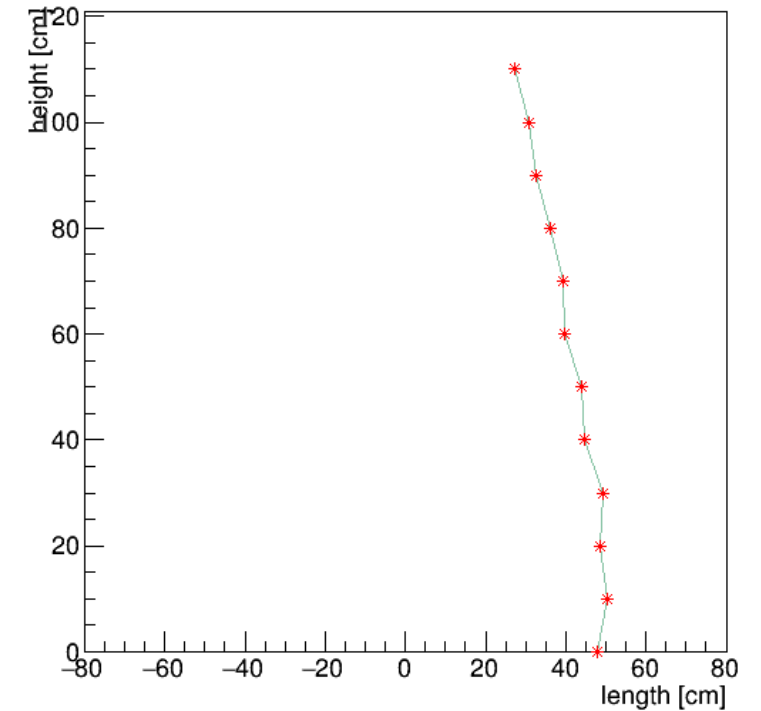
Before xCalibration (event=7)



After xCalibration (event=7)



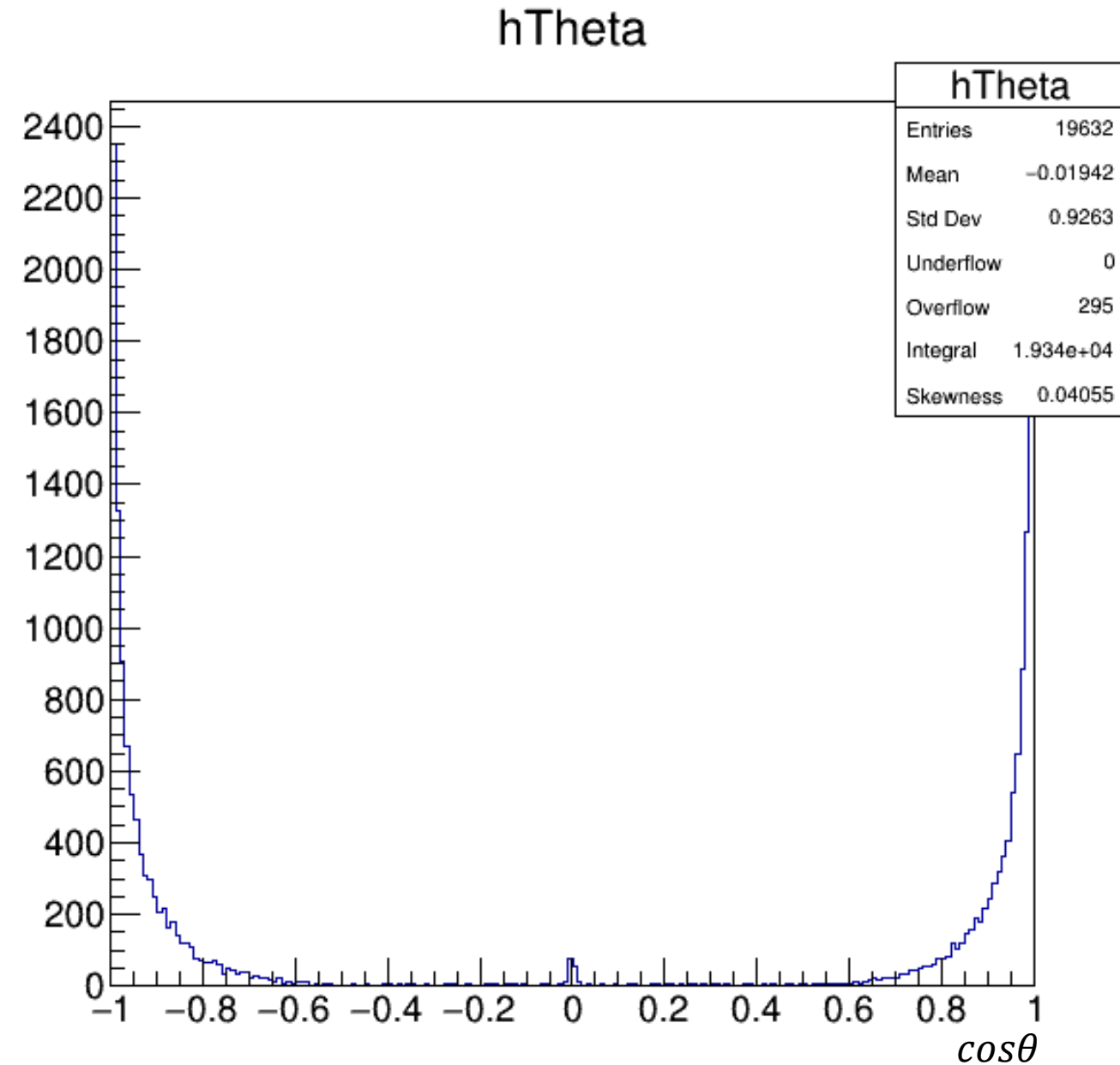
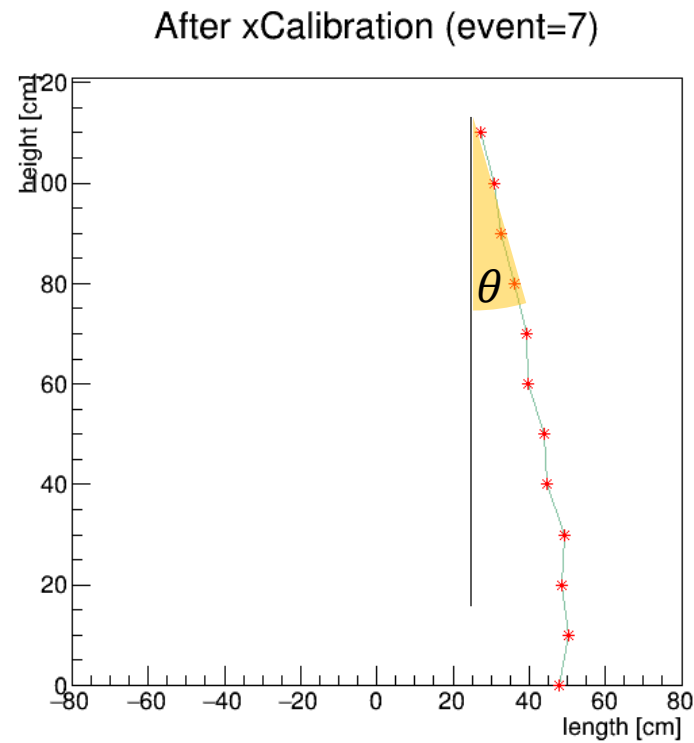
After xCalibration (event=7)



Performance

2D trajectory reconstruction

Angle distribution

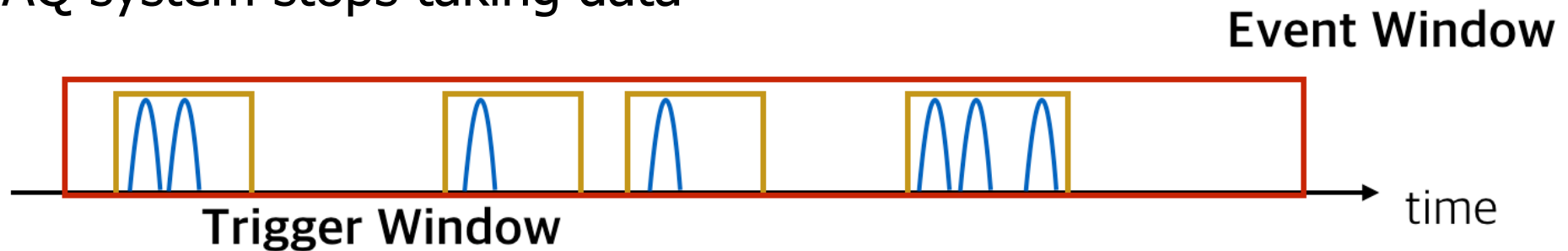


Simulation

Experimental procedure : Data taking

In the experiment, data will be taken in the following sequence.

- ① Data acquisition starts, by positron detachment laser(PDL) trigger, when an anti-hydrogen begins to fall
- ② If any signal hit(from annihilation or cosmic ray), data is taken with time stamp, which trigger window is about $1\mu\text{s}$
- ③ After the certain time of the event window(500ms), the DAQ system stops taking data



Simulation

Experimental procedure : Data analysis

For the single event,

- ① Clustering : merge hits of adjacent bars induced by a single track
- ② Cosmic-ray rejection : using the number of hits, hit pattern, and dT, determine if the trigger was caused by a cosmic ray or annihilation.

After that, the only real annihilation trigger should remain.

If so, we reconstruct the TOF, the annihilation vertex, and etc...

If not, the event may not be used

Good Events = Only one trigger selected

Successful Events = Real annihilation events in Good events

Cosmic-ray rejection

3 algorithms

- (1) Select the triggers having Top-Bottom hits(with two or more hits).
Compare **the earliest** top hit and **earliest** bottom hit.
And reject the trigger with $dT = T_1 - T_2 > 2$

T_1 = time to bottom of TOF, T_2 = time to top of TOF

- (2) Select the triggers having Top-Bottom hits(with two or more hits).
But compare **all the time differences** of top and bottom hit.
At least one combination of $dT > 2$ found, the event rejected.
- (3) Select the triggers having Top-Bottom-hits, **more than two hits**.
But compare **all the time differences** of top and bottom hit.
At least one combination of $dT > 2$ found, the event rejected.

Cosmic-ray rejection

Simulation results

Threshold : 0.3MeV, Event window : 500ms

	Algorithm 1	Algorithm 2	Algorithm 3
Selection Efficiency (Top / Bot. ann.)	0.4822 / 0.3458	0.4335 / 0.3389	0.4211 / 0.3342
Background Fraction (Top / Bot. ann.)	0.0212 / 0.0363	0.0242 / 0.0355	0.0152 / 0.0219

Selection efficiency = (the number of Good events) / (total events)

Selection accuracy = (the number of Successful events) / (Good events)

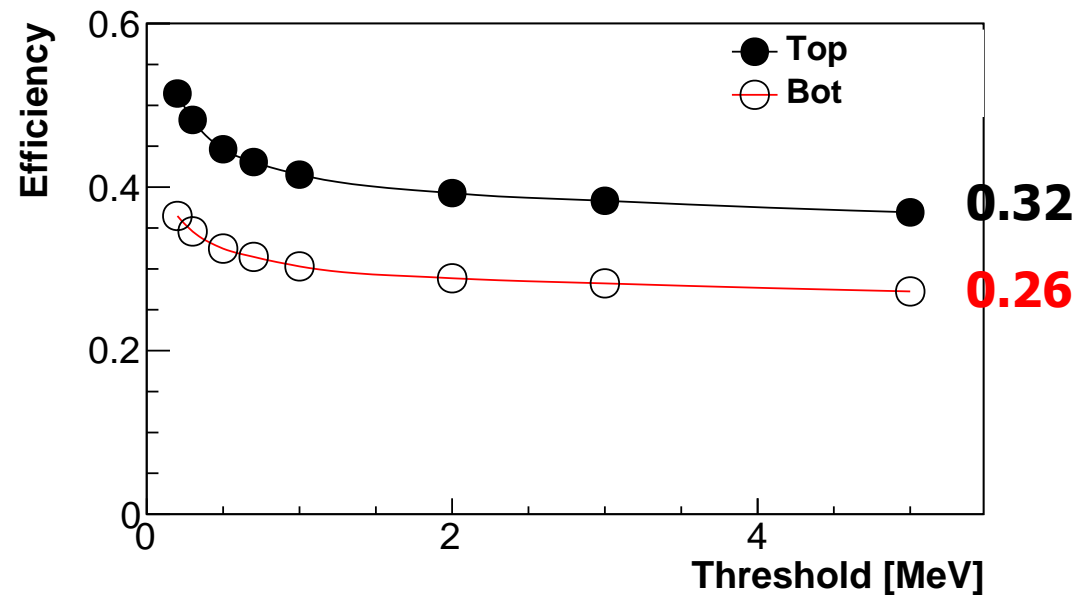
Background fraction = 1 – (Selection accuracy)

Cosmic-ray rejection

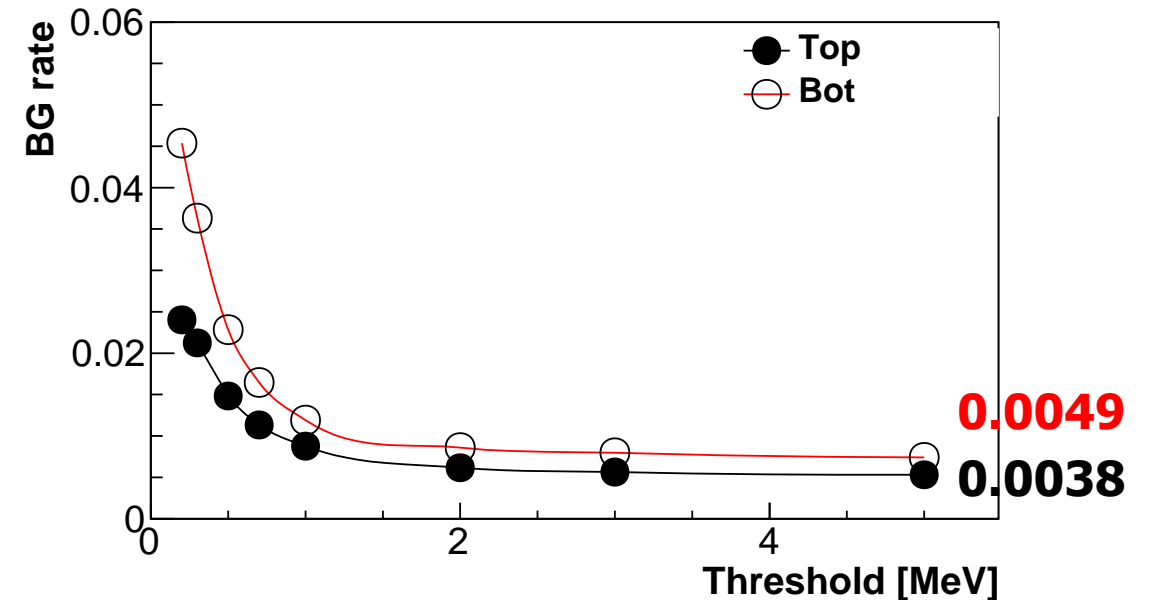
Simulation results – Algorithm 3

Threshold dependence (event window=500ms)

Cos. Rej. Eff. Thres. Dep. (EW = 500 ms)



Cos. Rej. BG Rate. Thres. Dep. (EW = 500 ms)

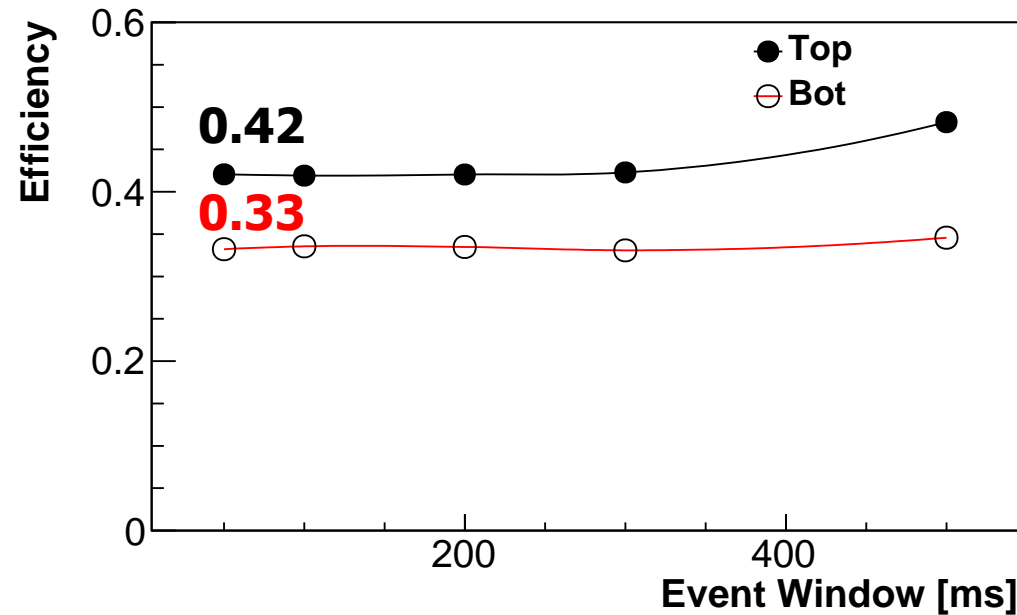


Cosmic-ray rejection

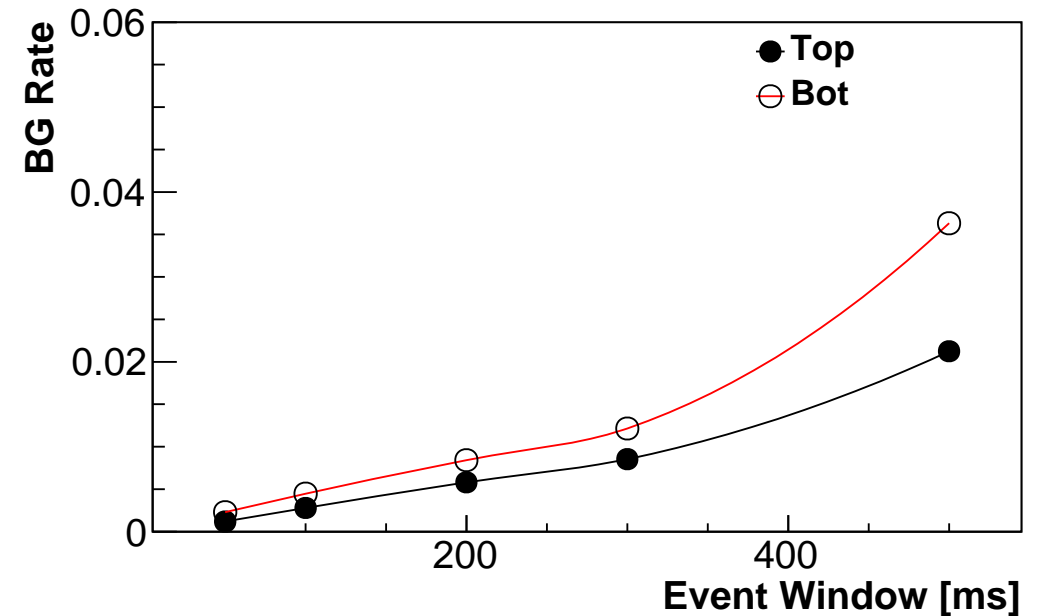
Simulation results – Algorithm 3

Event window dependence (th=0.3MeV)

Rej. Eff. Event Win. Dep. (Thres = 0.3 MeV)



BG Rate Event Win. Dep. (Thres = 0.3 MeV)

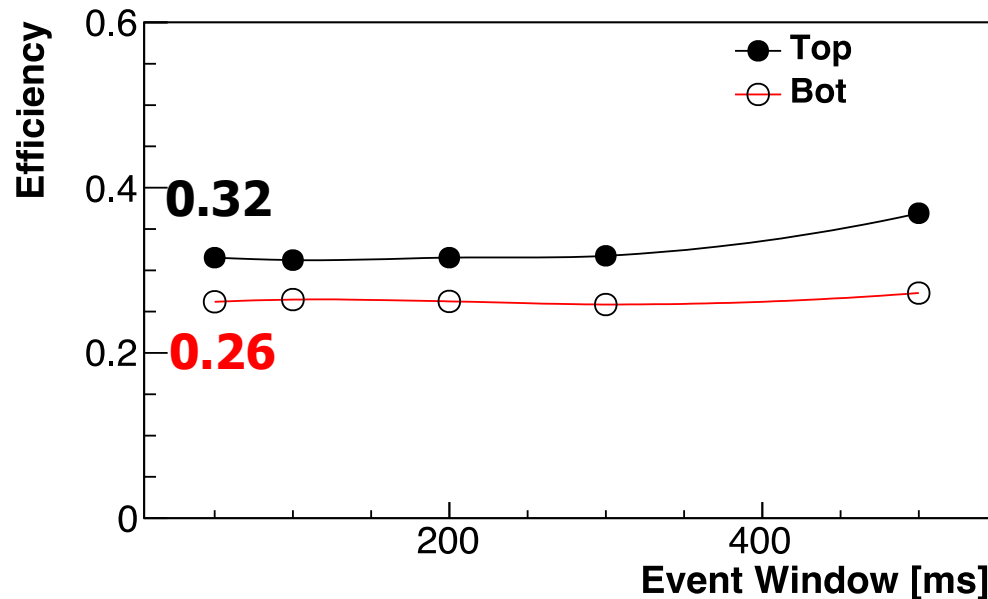


Cosmic-ray rejection

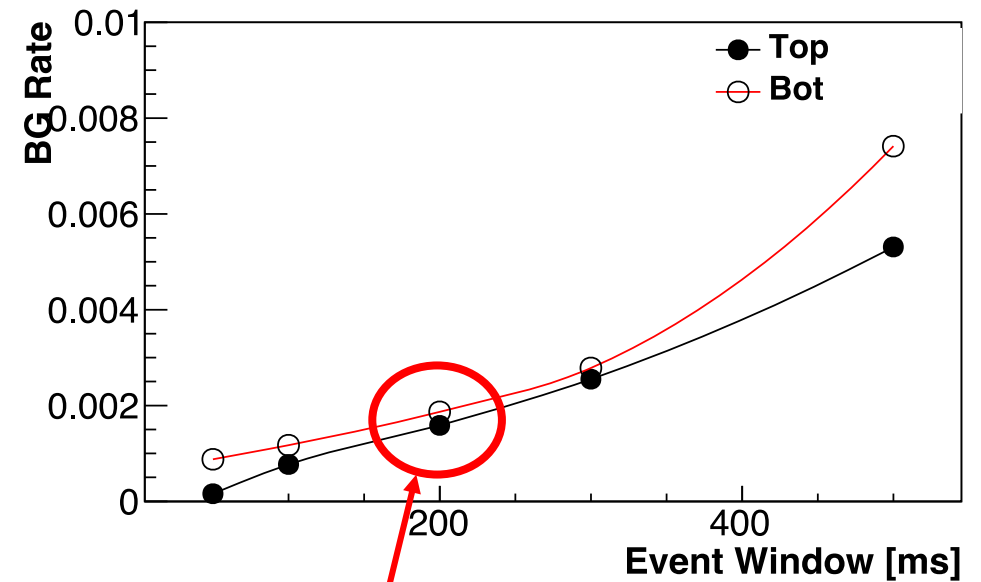
Simulation results – Algorithm 3

Event window dependence (th=5.0 MeV)

Rej. Eff. Event Win. Dep. (Thres = 5.0 MeV)



BG Rate Event Win. Dep. (Thres = 5.0 MeV)



About 0.2%

Summary

TOF will provide cosmic-ray rejection and free-fall time measurement.

Time resolution better than 0.2ns has been achieved, which is required for effective cosmic ray rejection and top/bottom annihilation separation.

With GEANT4 simulation, we checked performance of cosmic-ray rejection.

Efficiency and cosmic-ray background fraction for various cases has been studied

With high threshold and narrow event window, we can reduce background fraction below 0.2%

Material information of MMC, FFC and other needs to be added in the simulation.

* Note : Thicknesses of flange are 6mm(bottom, side) and 30mm(top)