Review on Proposed antimatter gravity measurement with an antihydrogen beam

A. Kellerbauer, Nuclear Instruments and Methods in Physics Research B 266(2008)

Aram lee

- 1. Motivation
- 2. AEGIS experiments
 - a. Aims
 - b. Methods
- 3. Conclusion

1. Motivation



• Antimatter : a material composed of anti-particles



Marco G., AEGIS at CERN : measuring Antihydrogen fall, INFN, Posmol(2011)

1928, prediction of antimatter (Dirac)
1932, discovery of positron in cosmic rays (Anderson)
1955, antiproton discovery (Segre')
1956, antineutron discovery (Cork)
1995, creation of high-energy antihydrogen (CERN)
2002, creation of 10K antihydrogen (ATHENA, ATRAP)
2011, antihydrogen confinement (ALPHA)

Weak equivalence principle(WEP), in the general relativity

" In a uniform gravitational field, all objects fall with the same acceleration, regardless of their composition."

WEP is tested with matter to high accuracy : $\sim 10^{-16}$ (2016)

But the behavior of anti-matter in a gravitational field has never been tested experimentally.



Weak equivalence principle(WEP), in the general relativity

Gravity is the only force not described by a quantum field theory.

According to a hypothetical quantum theory of gravity, gravity could have with a tensor, vector, and scalar component. A vector component might lead a repulsive force. That constitute a violation of the WEP.

Anti-gravity can be a possible solution for cosmological problem. But still a number of arguments against.

In order to repair contradictions and to incorporate gravity into the Standard model, observing the behavior of antimatter in gravitational field is necessary.



1. Motivation

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the AEGIS(Antimatter Experiment : Gravity, Interferometry, Spectroscopy) experiments

: the direct measurement of the Earth's local gravitational acceleration \bar{g} on \bar{H} with 1% precision



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① Ps production and excitation

A sizable fraction of Ps diffuses out of the nanoporous material film.



The yield depend on the target material, temperature and the implantation depth. (~30% in silicon-based polymer materials cooled to 50K)

① Ps production and excitation

The cross-section of the \overline{H} formation reaction : $\sigma \approx a_0 n^4 \rightarrow \text{excitation of Ps is needed}$

The Ps excited to a Rydberg state with principal quantum number n=30...40by two pulsed-laser systems.



(2) \overline{H} beam formation

The charge exchange reaction

 $Ps^* + \bar{p} \rightarrow \bar{H}^* + e^-$ (* denotes a highly excited Rydberg state.)

- large cross-section ($\approx 10^7 \text{\AA}^2$ for n = 35)
- narrow and well-defined band of \overline{H}^*
- determination of velocity distribution of \overline{H}^* by \overline{p} temperature (~25 80m/s)

(2) \overline{H} beam formation

Stark acceleration

Atoms are accelerated or decelerated (experienced the force) by electric field gradient. Rydberg atoms which have large dipole moment are amenable. they are accelerated into a beam by inhomogeneous electric field.





< Principle sketch until now >

- 1. e^+ capture
- 2. \bar{p} production
- 3. Ps formation
- 4. Ps excitation
- 5. \overline{H} production
- *6.* \overline{H} beam formation

Note : \bar{p} trap before acceleration

Gravity measurement 3

Free-fall of \overline{H} cannot be detected in usual way



 $h \sim 20 \mu m$, where $L \sim 1m$, $v_x \sim 500m/s$ much larger than beam radial size after 1m flight (~ cm)

③ Gravity measurement

Free-fall detection by Moiré deflectometer





S. Aghion, A moiré deflectometer for antimatter, nature communications(2014)

Free-fall detection by Moiré deflectometer



Grating transparency $\sim 30\%$

a : the grating period (~80μm) d : the gap in gratings

L : the length between gratings($\sim 40cm$)

It is NOT a quantum deflectometer. No diffraction occurs.



③ Gravity measurement

Free-fall detection by Moiré deflectometer



$$\delta = 2 * \left(\frac{1}{2}gT^2\right) / a = \frac{gT^2}{a}$$

Extraction of g value from observables(T, δ) – illustrated by Monte-carlo simulations

First, record vertical position for each event as a function of T(or v)



And bin these events in symmetric classes of T² $v_{beam} = 600 - 550 - 500 - \dots - 200 \text{ m/s}$

Extraction of g value from observables(T, δ) – illustrated by Monte-carlo simulations

Next, extract δx of the fringe pattern for each of the count classes



Extraction of g value from observables(T, δ) – illustrated by Monte-carlo simulations

Finally, plot δx against the mean time of flight $\sqrt{\langle T^2 \rangle}$



Extraction of g value from observables(T, δ) – illustrated by Monte-carlo simulations

Finally, plot δx against the mean time of flight $\sqrt{\langle T^2 \rangle}$

the coefficient of a quadratic fit to that yields \bar{g}

$$\delta x = \delta * a = gT^2$$

(1% relative precision, about $10^5 \overline{H}$ atoms at 100mK)



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| | | | ATHENA | | Charge exchange production of \overline{H} |
|---|-------|---|-------------|---|--|
| • | AEGIS | = | ATRAP | + | Stark acceleration |
| | | | (at the AD) | | The Moiré deflectometer |

- Obtaining \overline{H} atoms at 100mK is an essential part, since the precision of gravity measurement is mainly limited by the \overline{H} temperature.
- If even colder \overline{H} ensembles can be attained, gravity measurements with even higher precision as well as competitive CPT tests through spectroscopy.